

# Seismic Analysis of the Pulsating Subdwarf B Star EPIC 212508753 Using Data From NASA's Kepler Space Telescope

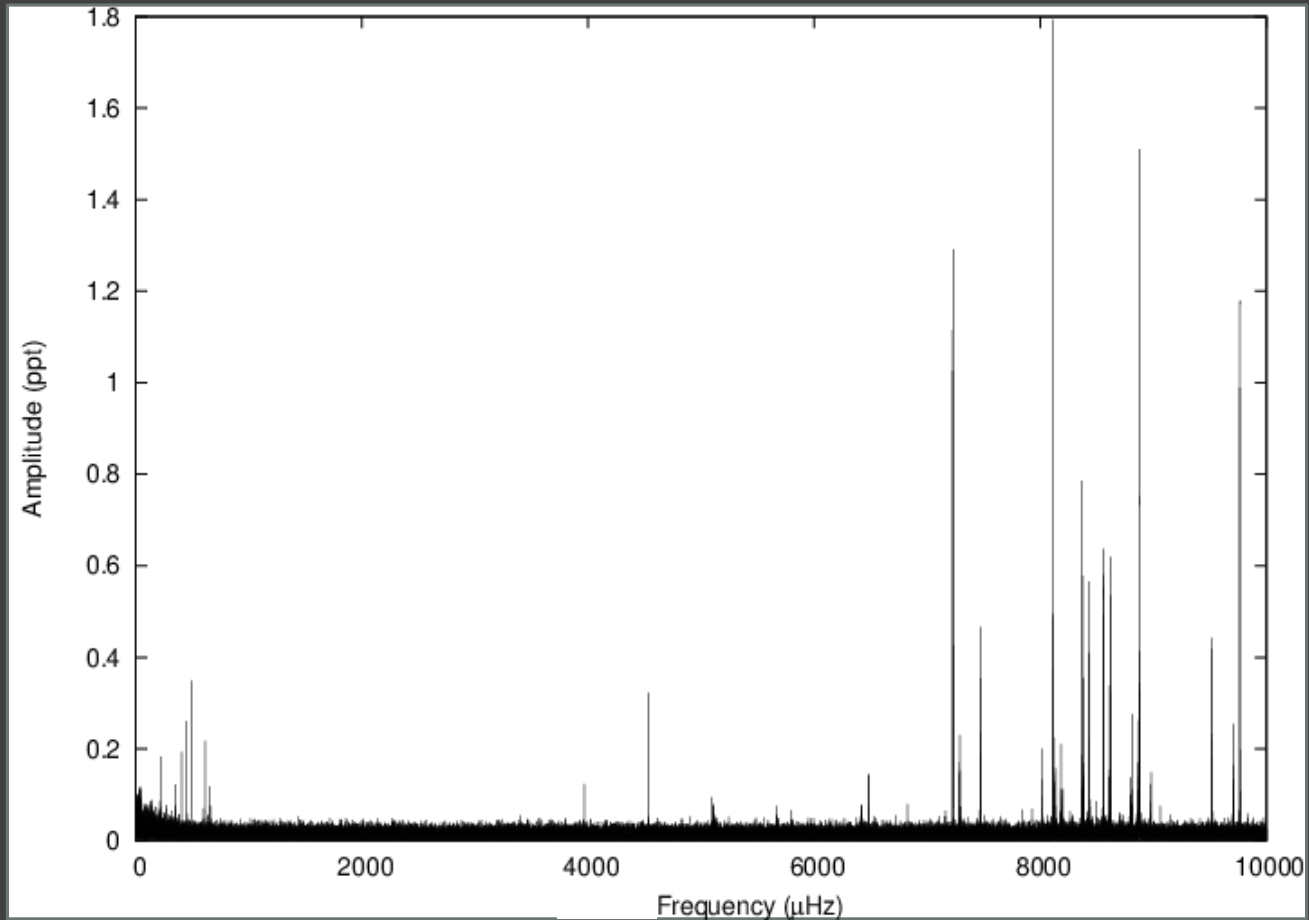
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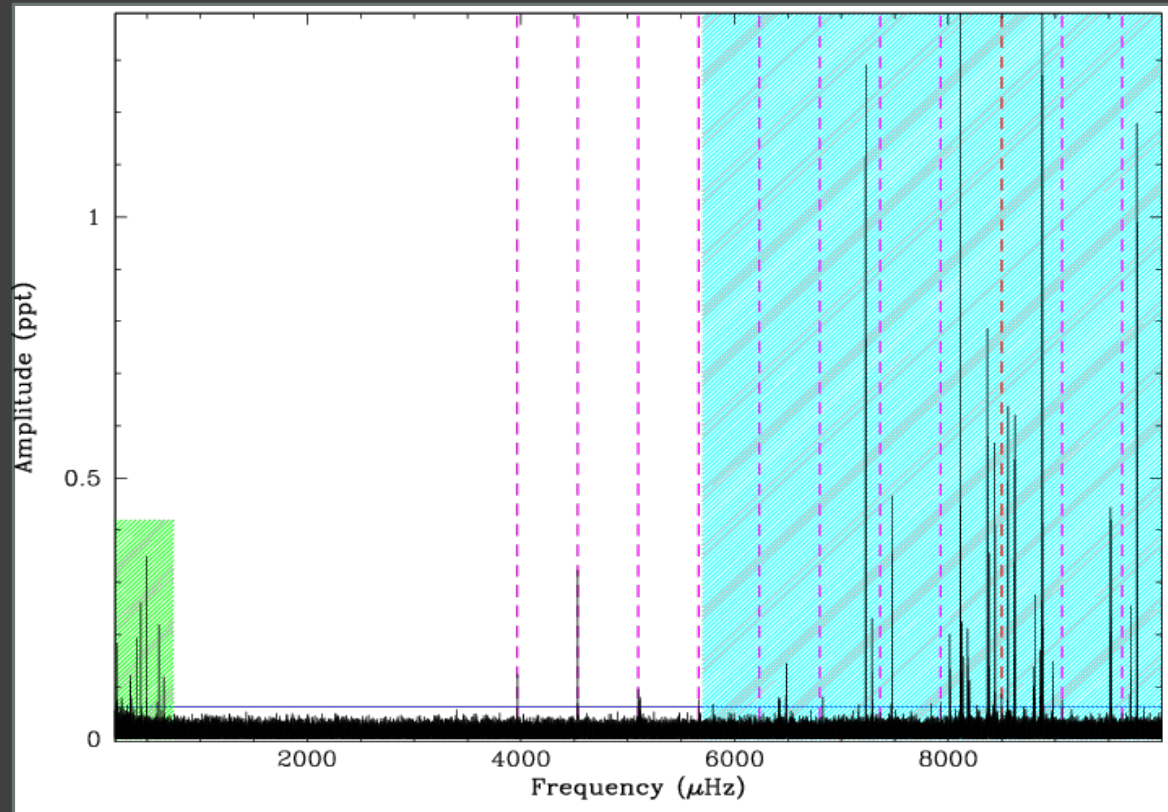
# Fourier Transform (FT)

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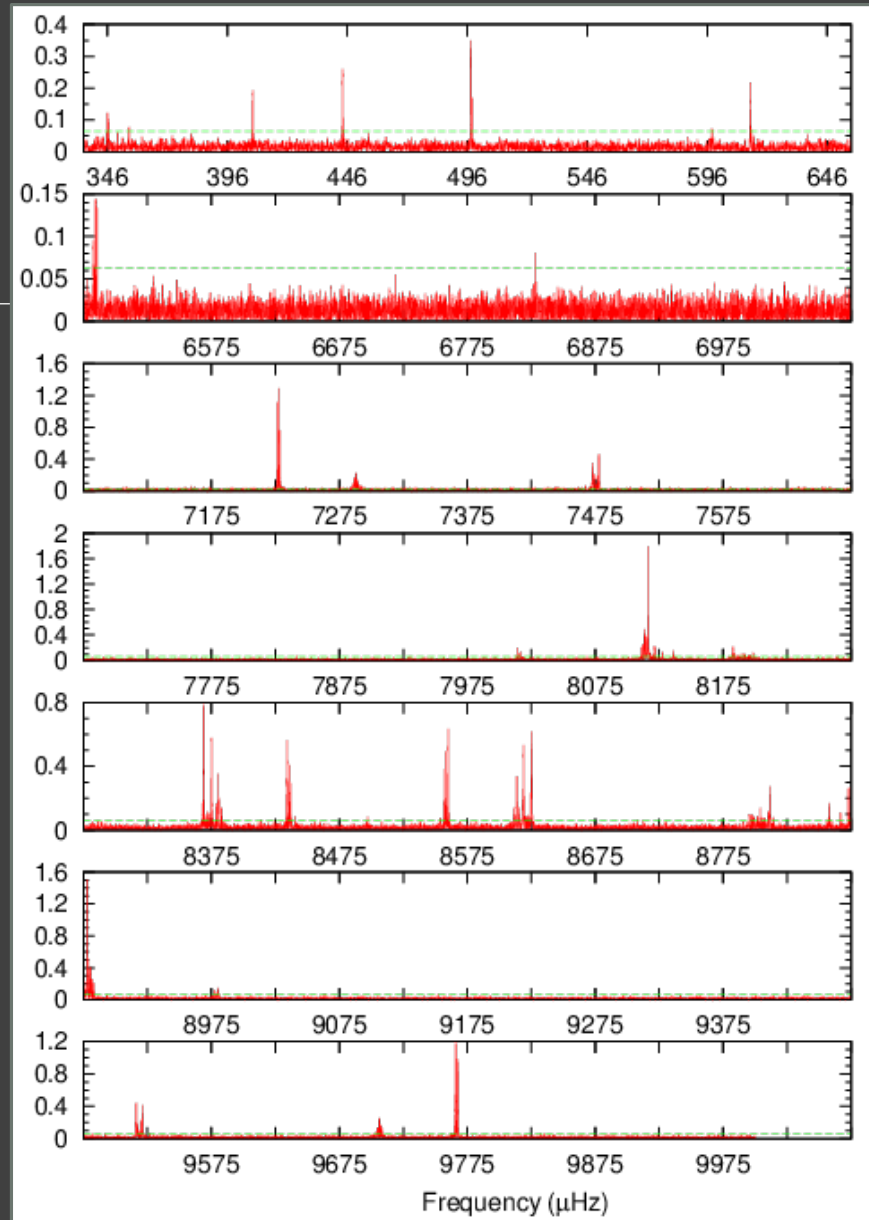
# Fourier Transform (FT)

- Resolution of  $0.147 \mu\text{Hz}$
- Green = g-modes
- Blue = p-modes
- Purple = artifacts
- Red = Nyquist



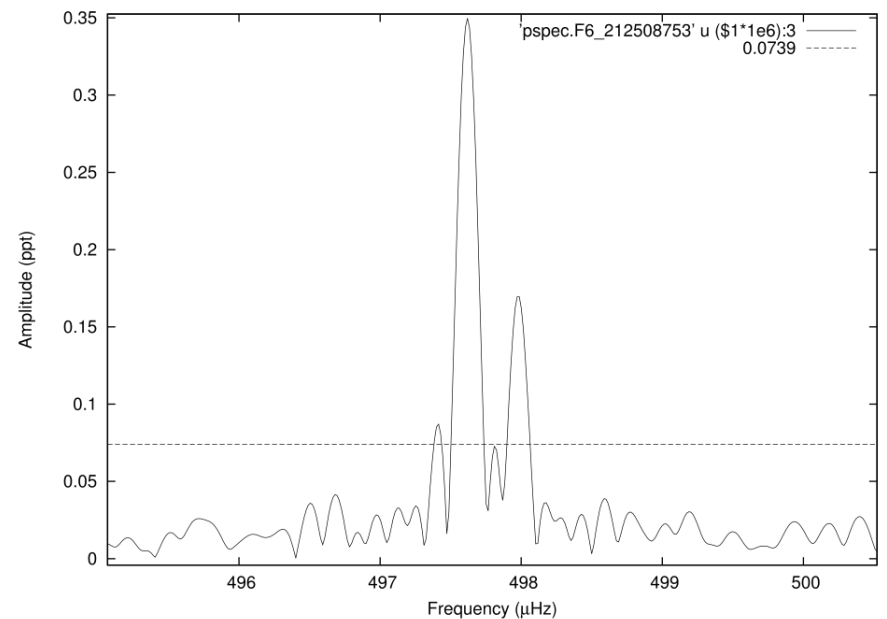
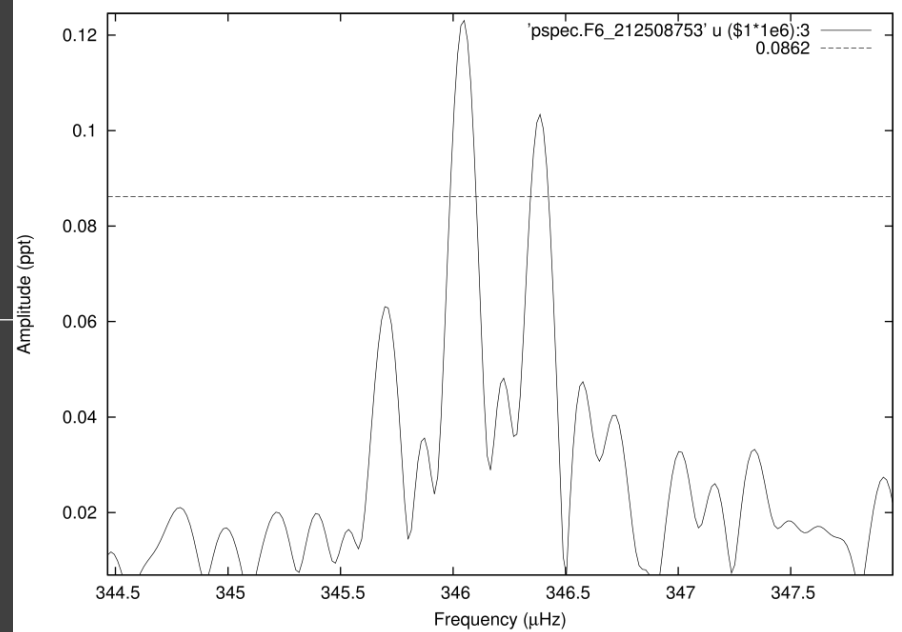
# Analysis

- Pulsations above or near the detection threshold were examined
  - Known artifacts were ignored



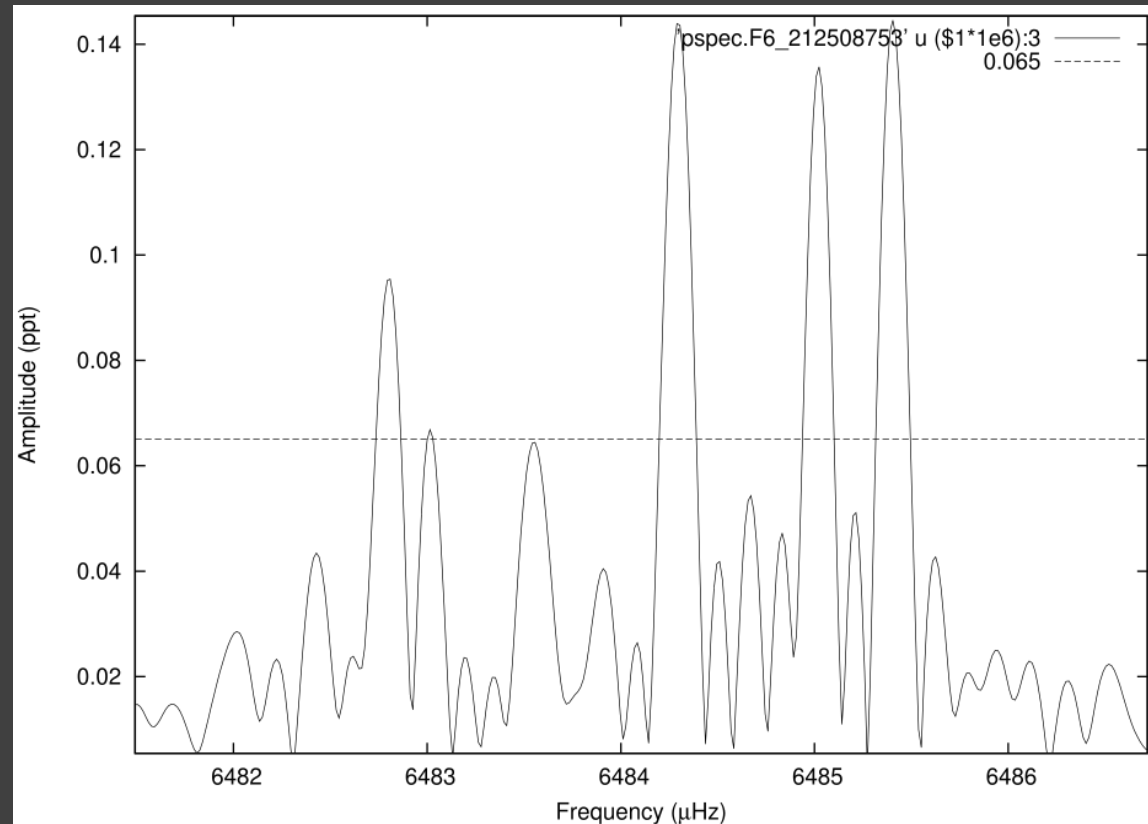
# g-modes

- 8  $n$  values found
- Average multiplet splitting of  $0.360 \mu\text{Hz}$
- Average period spacing of  $227.7 \pm 1.0 \text{ s}$ 
  - Short side of average



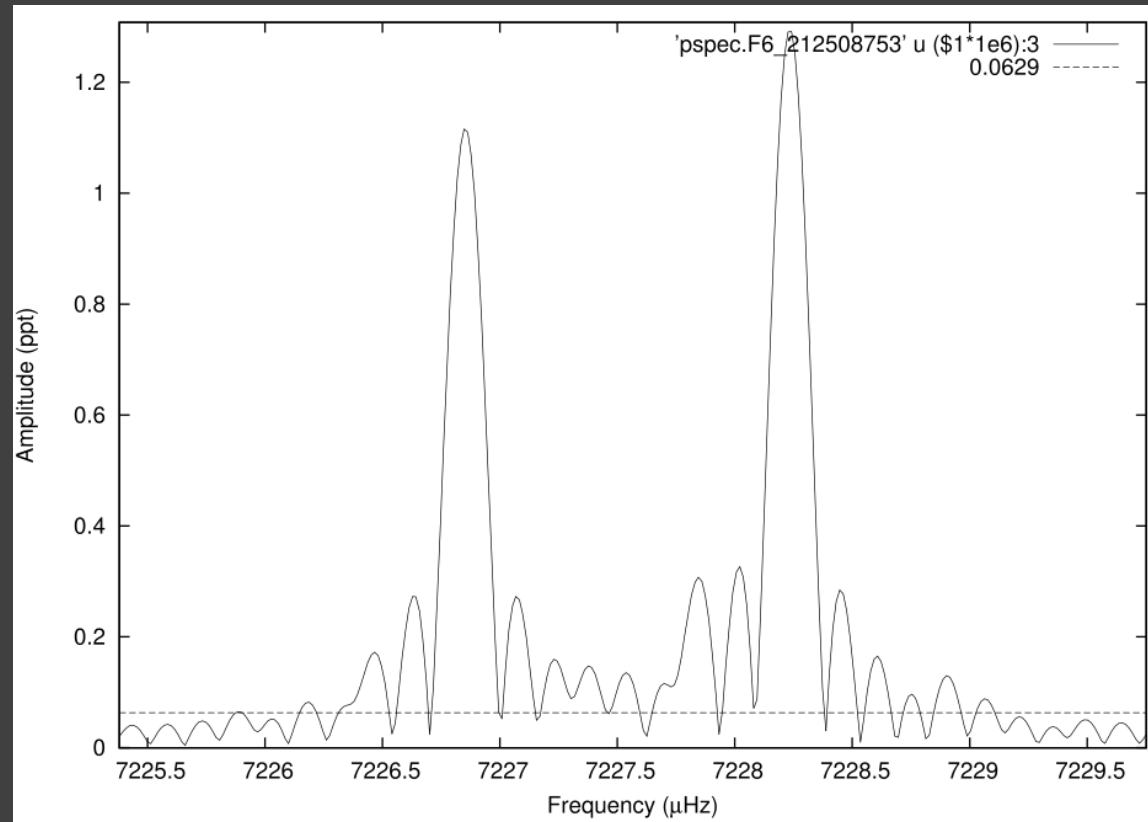
# p-modes

- 13 n values identified
- Multiplet splitting of  $0.721 \mu\text{Hz}$
- Example of  $l=2$



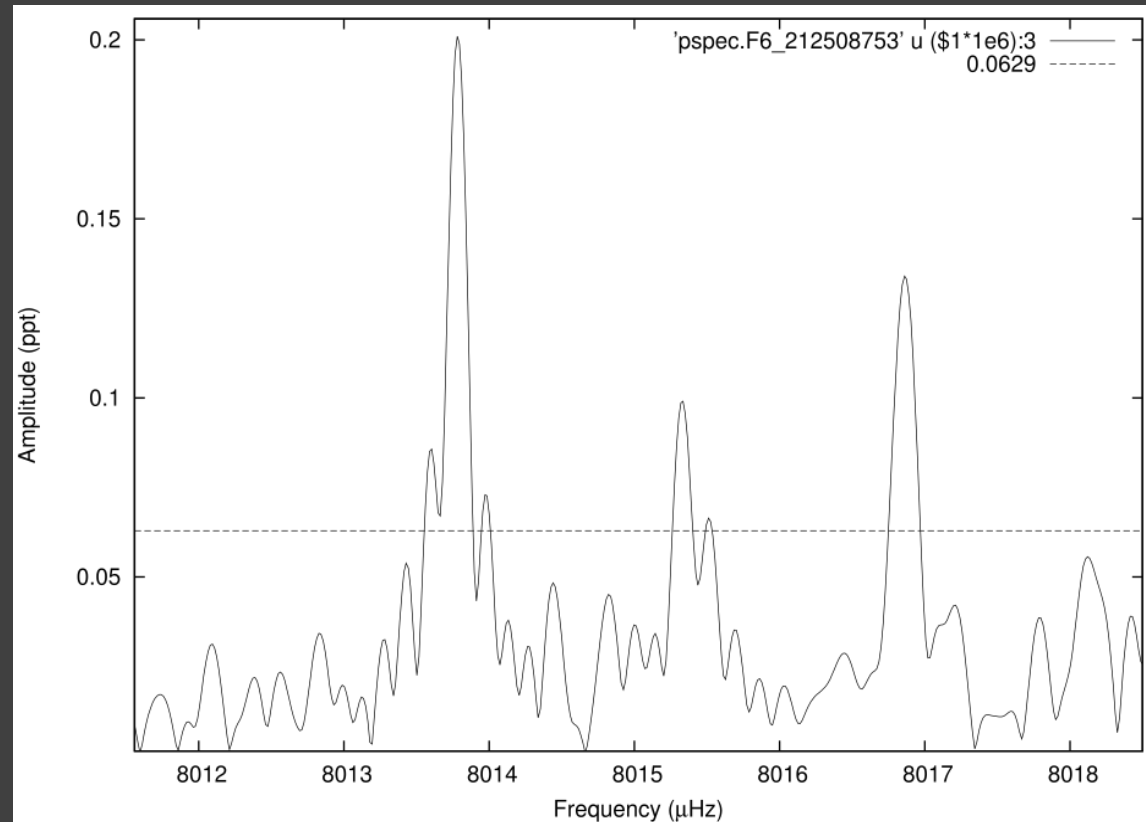
# p-modes

- 13 n values identified
- Multiplet splitting of  $0.721 \mu\text{Hz}$
- Example of  $l=1$



# p-modes

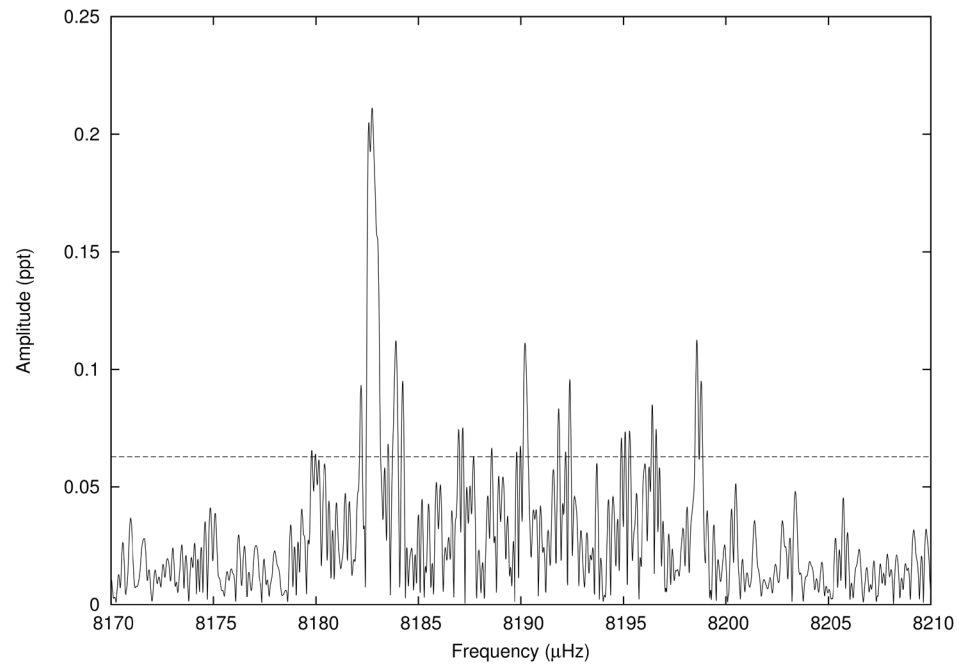
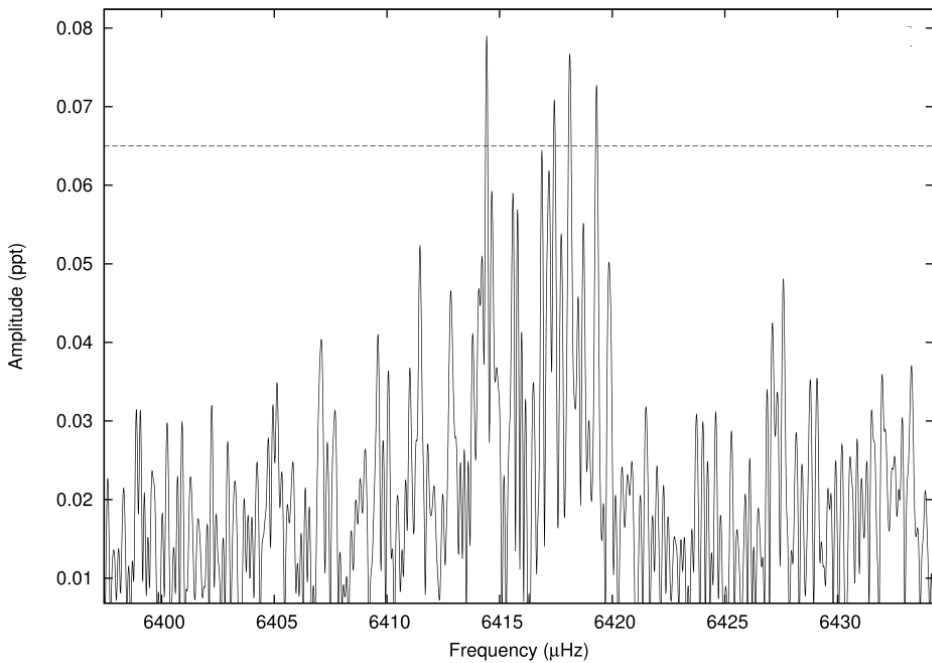
- 13 n values identified
- Multiplet splitting of  $0.721 \mu\text{Hz}$
- Example of  $l=2$





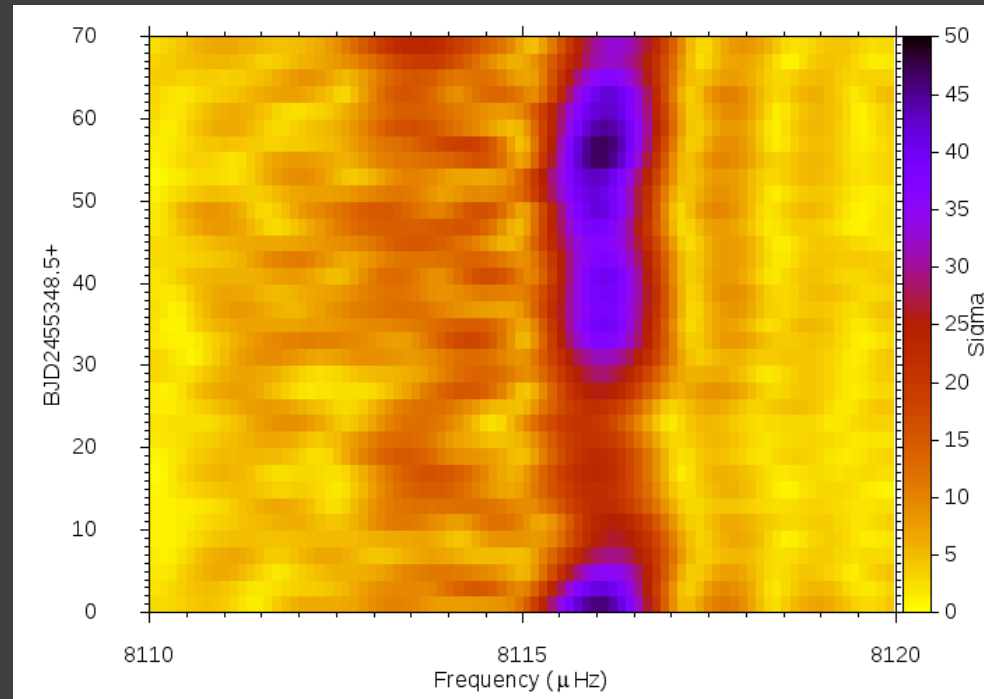
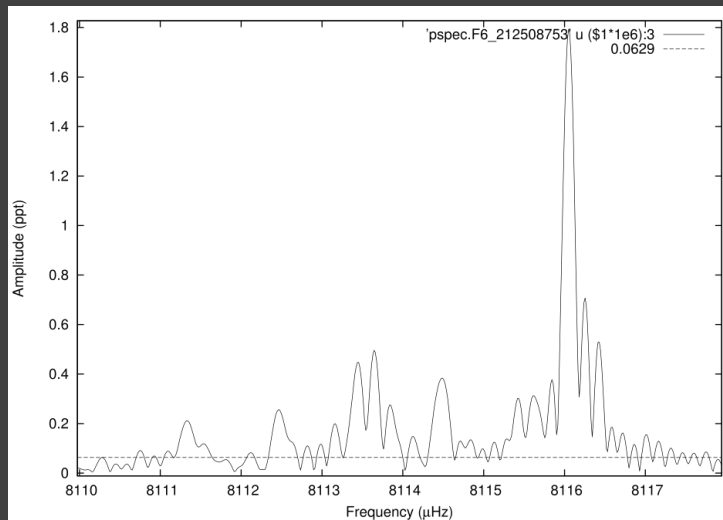
# Stochastic Oscillations

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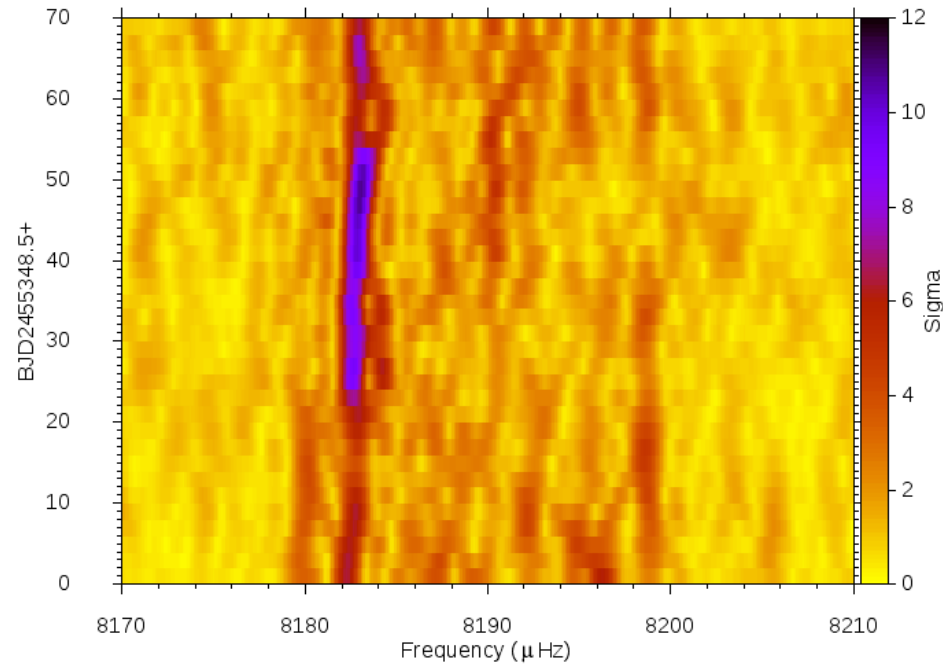
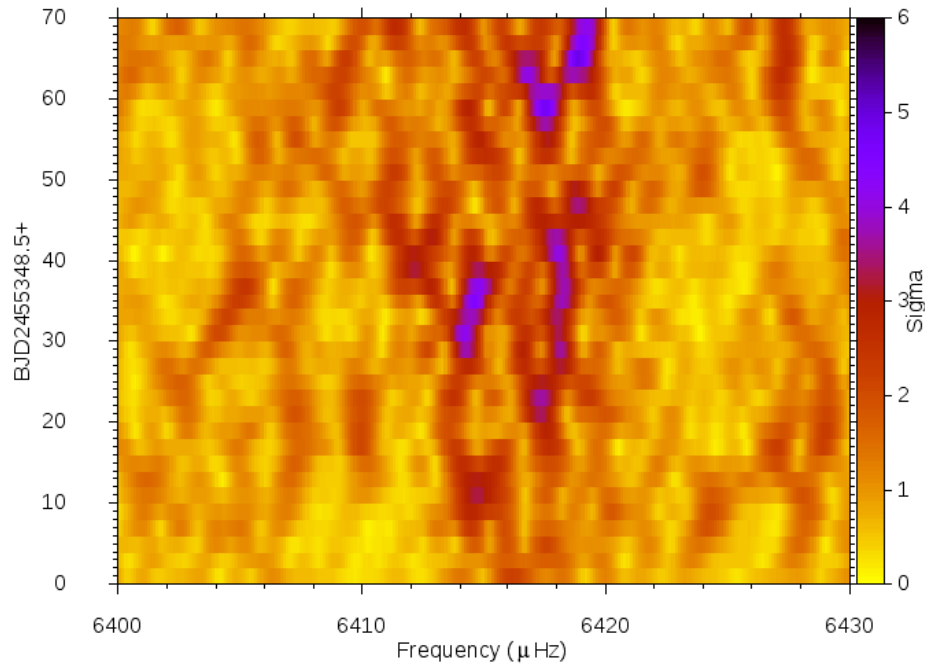
# Sliding Fourier Transforms

- Used to examine change in pulsations over time
- 10 day bins
- 2 day steps



# Stochastic Oscillations

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# Results

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- Ledoux Rotational Splitting Relationship:  $\nu_{n,l,m} = \nu_{n,l,0} + m\Omega(1 - C_{n,l})$ 
  - $\Omega$ : Rotation Period
  - $C_{n,l}$ : Ledoux Constant
- g-mode region
  - $C_{n,l} = \frac{1}{l(l+1)} = \frac{1}{2}$
  - $\Delta\nu = 0.360 \mu\text{Hz}$  for  $l, m = 1$
  - $\Omega \approx 16.1 \text{ days}$
- p-mode region
  - $C_{n,l} \approx 0$
  - $\Delta\nu = 0.721 \mu\text{Hz}$  for  $m = 1$
  - $\Omega \approx 16.1 \text{ days}$
- Solid-body rotation!

# Results

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- Spectral analysis done by John Telting of the Nordic Optical Telescope

	sdB Component	Main Sequence Companion Component
$T_{eff}$	$36,320 \pm_{710}^{500} K$	$6,457 \pm_{560}^{460} K$
$\log(g)$	$5.62 \pm 0.09$	$4.419 \pm_{0.2}^{0.1}$
$\log\left(\frac{nHe}{nH}\right)$	$-0.98 \pm_{0.15}^{0.03}$	

- Hottest Kepler-observed sdB pulsator
- Should not have g-modes; models can't drive them past 24,000 K
- F7V companion star

# Conclusion

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- Hybrid pulsator with most power in the p-mode region
- Very unique Kepler-observed sdBV
  - Hottest!
  - Still has g-modes!
  - Shortest pulsation periods!
  - Stochastic oscillations!
  - Solid-body rotator!
  - Relatively short rotation period!

